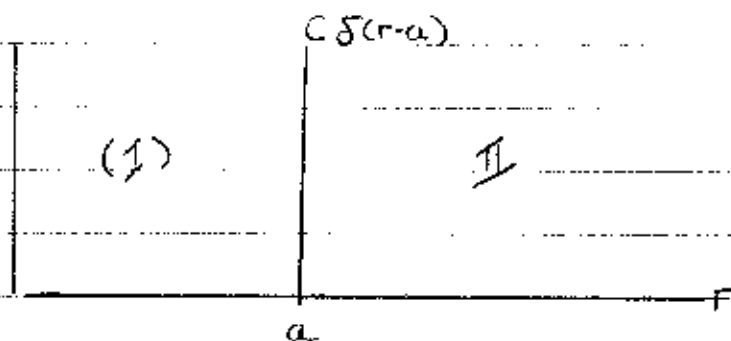


Home work 7, solutions
problem 1, solution.

We use the partial wave method. Since $ka \ll 1$, only s-wave scattering is important.

We need to find the phase shift δ_0 for $V(r) = C \delta(r-a)$.



Region I: $U(r) = A \sin kr$, $k^2 = \frac{2mE}{\hbar^2}$.

Region II: $U(r) = B \sin(kr + \delta_0)$.

Boundary conditions: $U(r)$ is continuous at $r=a$, but

$\frac{dU}{dr}$ is not, $\frac{d^2U}{dr^2} + \frac{2m}{\hbar^2} (E - V(r))U = 0$.

$$\left[\frac{dU_2}{dr} \Big|_{a+\epsilon} - \frac{dU_1}{dr} \Big|_{a-\epsilon} \right]_{\epsilon \rightarrow 0} = \left[-\frac{2m}{\hbar^2} \int_{a-\epsilon}^{a+\epsilon} (E - C\delta(r-a))U(r) dr \right]_{\epsilon \rightarrow 0} = \frac{2m}{\hbar^2} C U(a).$$

We therefore have $A \sin ka = B \sin(ka + \delta_0)$,

$kA \cos ka = kB \cos(ka + \delta_0) - \frac{2m}{\hbar^2} CA \sin ka$.

$k \cot ka + \frac{2m}{\hbar^2} C = k \cot(ka + \delta_0)$.

$\delta_0 = \cot^{-1} \left(\cot ka + \frac{2m}{\hbar^2} \frac{C}{k} \right) - ka$, $G_k = \frac{4\pi}{k^2} \sin^2 \delta_0$.

problem 2, solution

As $E \rightarrow 0$, $k \rightarrow 0$, $ka \rightarrow 0$. If $ka \ll 1$ then the scattering is dominated by s-wave scattering.

$$\text{Then } \delta_k(\theta) = \frac{1}{k^2} \sin^2 \delta_0(k), \quad \delta_k = \frac{4\pi}{k^2} \sin^2 \delta_0.$$

However, when $\delta_0 = \pi$, $\delta_k(\theta)$ and δ_k are ≈ 0 . This is the Ramsauer-Townsend effect. In this case the $l=1$ partial wave contribution to the scattering is the dominant contribution and

$$\delta_k(\theta) = \frac{1}{k^2} 9 \sin^2 \delta_1 \cos^2 \theta, \quad \delta_k = \frac{4\pi}{k^2} 3 \sin^2 \delta_1.$$

To find what $V_0 a^2$ has to be for $\delta_0 = \pi$, we need to find the expression for δ_0 for a 3-dimensional square well potential.

$$V(r) = -V_0 \quad r < a, \quad U_{k_0}(r) = C_1 \sin \sqrt{k^2 + U_0} r, \quad k^2 = \frac{2mE}{\hbar^2}, \quad U_0 = \frac{2mV_0}{\hbar^2} \\ V(r) = 0 \quad r > a, \quad U_{k_0}(r) = C_2 \sin(kr + \delta_0).$$

$U_{k_0}(r)$ and $\frac{d}{dr} U_{k_0}(r)$ are continuous at $r=a$. This implies $\tan(ka + \delta_0) = \frac{k}{\sqrt{k^2 + U_0}} \tan \sqrt{k^2 + U_0} a$.

$$\text{We want } \delta_0 = \pi, \quad \tan ka = \frac{k}{\sqrt{k^2 + U_0}} \tan \sqrt{k^2 + U_0} a.$$

$$\text{As } ka \rightarrow 0 \text{ this becomes } ka = \frac{k}{\sqrt{U_0}} \tan \sqrt{U_0} a.$$

$$\sqrt{U_0} a = \tan \sqrt{U_0} a, \quad \sqrt{U_0} a = n\pi + b.$$

$$n=0 \quad b=0.$$

$$n=1 \quad b=1.352. \quad \underline{\sqrt{U_0} a = 4.494} \quad \underline{\tan \sqrt{U_0} a = 4.497}.$$

problem 3, solution

$$a) \left. \frac{d\sigma^B}{d\Omega} \right|_{\vec{k}_f, \vec{k}_i} = \frac{m^2}{4\pi^2 \hbar^4} \left| \int d^3r e^{i\vec{q}\cdot\vec{r}} V(\vec{r}) \right|^2 \quad \text{with } \vec{q} = \vec{k}_i - \vec{k}_f \text{ and } |\vec{k}_i| = |\vec{k}_f|.$$

$$V(\vec{r}) = V_0 e^{-\alpha r} \frac{1}{r}, \quad V_0 = -Ze^2, \quad \alpha = \frac{1}{a}.$$

$$\int d^3r e^{i\vec{q}\cdot\vec{r}} \frac{e^{-\alpha r}}{r} = \frac{4\pi}{q} \frac{q}{\alpha^2 + q^2}.$$

$$\left. \frac{d\sigma^B}{d\Omega} \right|_{\vec{k}_f, \vec{k}_i} = \frac{4m^2 V_0^2}{\hbar^4} \frac{1}{(\alpha^2 + q^2)^2}.$$

With $q = 2k \sin \frac{\theta}{2}$ we have $\left. \frac{d\sigma^B}{d\Omega} \right|_{\vec{k}_f, \vec{k}_i} = \frac{4m^2 Z^2 e^4 / \hbar^4}{\left(\frac{1}{a^2} + 4k^2 \sin^2 \frac{\theta}{2}\right)^2}$

b) Let $a \rightarrow \infty$, $\alpha \rightarrow 0$. Then $\left. \frac{d\sigma^B}{d\Omega} \right|_{\vec{k}_f, \vec{k}_i} = \frac{4m^2 Z^2 e^4}{\hbar^4} \frac{1}{16 \sin^4 \frac{\theta}{2} k^4}$.

This is the Rutherford cross section.

c) ratio: $\frac{\left. \frac{d\sigma}{d\Omega} \right|_a}{\left. \frac{d\sigma}{d\Omega} \right|_b} = \left(\frac{q^2}{\frac{1}{a^2} + q^2} \right)^2 = \left(\frac{a^2 q^2}{1 + a^2 q^2} \right)^2 = \left(\frac{\chi^2}{1 + \chi^2} \right)^2, \quad \chi = aq.$

As $\chi \rightarrow 0$ ratio $\rightarrow 0$, $\chi \rightarrow 0 \Rightarrow a \rightarrow 0 \Rightarrow \alpha \rightarrow \infty$.

The nuclear charge is completely screened. We have a neutral object.

As $\chi \rightarrow \infty$ ratio $\rightarrow 1$, $\chi \rightarrow \infty \Rightarrow a \rightarrow \infty \Rightarrow \alpha \rightarrow 0$

We have a bare nucleus, the screening vanishes.

problem 4, solution

$$a) 4\pi \int_0^{\infty} \rho(r) r^2 dr = \alpha 4\pi \int_0^{\infty} r^2 e^{-\beta r} dr = 4\pi \frac{\alpha}{\beta^3} \int_0^{\infty} r^{1\alpha} e^{-r^1} dr^1 = \frac{8\pi\alpha}{\beta^3}$$

$$1.6 \times 10^{-19} \text{ C} = 4.8 \times 10^{-10} \text{ esu} = \frac{8\pi\alpha}{\beta^3}$$

$$b) \text{ Assume } |\psi(r^2)|^2 \propto \rho(r), \quad 4\pi \int_0^{\infty} |\psi(r)|^2 r^2 dr = 1$$

$$|\psi(r)|^2 = \frac{\rho(r) \beta^3}{8\pi\alpha} = \frac{\beta^3 e^{-\alpha r}}{8\pi}$$

$$\langle r^2 \rangle = \int r^2 |\psi(r)|^2 dr = 4\pi \int_0^{\infty} r^4 |\psi(r)|^2 dr = \frac{1}{2} \beta^3 \int_0^{\infty} r^4 e^{-\beta r} dr$$

$$= \frac{1}{2\beta^2} \int_0^{\infty} r^{14} e^{-r^1} dr^1 = \frac{4!}{2\beta^2} = \frac{12}{\beta^2}$$

$$c) \text{ Assume } \sqrt{\langle r^2 \rangle} = 10^{-15} \text{ m} = 10^{-13} \text{ cm}, \quad \beta^2 = \frac{12}{\langle r^2 \rangle} = 1.2 \times 10^{27} \text{ cm}^{-2}$$

$$\alpha = 4.8 \times 10^{-10} \text{ esu} \frac{\beta^3}{8\pi} = 7.9 \times 10^{29} \text{ esu cm}^{-3}$$

problem 5, solution

In the Born approximation we have

$$\sigma_k^B(\theta, \phi) = \frac{m^2}{4\pi^2 \hbar^4} \left| \int d^3r' e^{-i\vec{q} \cdot \vec{r}'} V(\vec{r}') \right|^2, \quad \vec{q} = \vec{k} - \vec{k}', \quad q = 2k \sin \frac{\theta}{2}.$$

$$B \int d^3r' e^{-i\vec{q} \cdot \vec{r}'} \delta(\vec{r}') = B.$$

$$\sigma_k^B(\theta, \phi) = \frac{m^2 B^2}{4\pi^2 \hbar^4}.$$

The differential scattering cross section is independent of scattering angle and velocity.

$$\sigma_k^B(\theta, \phi) = \frac{m^2 B^2}{4\pi^2 \hbar^4}.$$